

Equilibrium Climate

Jintai Lin

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思考题

Derive a relationship for the height of a given pressure surface (p) in terms of the pressure p_0 and temperature T_0 at sea level, assuming that the temperature decreases uniformly with height at a rate of Γ K km⁻¹

$$z = \frac{T_0}{\Gamma} \left[1 - \left(\frac{p}{p_0} \right)^{R\Gamma/g} \right]$$

高度计

This is the basis for the calibration of aircraft altimeters

思考题

Lifting all assumptions for air parcel, except that the environment is still in hydrostatic equilibrium.

- (a) Show that when a parcel of dry air at temperature T' moves adiabatically in ambient air with temperature T , the temperature lapse rate of the air parcel is given by

$$\Gamma = -\frac{\partial T'}{\partial z} = \frac{T'}{T} \frac{g}{c_p}$$

- (b) Explain why the lapse rate in this case differs from the dry adiabatic lapse rate (g/c_p)

思考题

- Assuming the truth of the second law of thermodynamics, prove that an isolated ideal gas can expand spontaneously (e.g., into a vacuum) but cannot contract spontaneously
- One kilogram of ice at 0°C is placed in an isolated container with 1 kg of water at 10°C and 1 atm. (a) How much of the ice melts? (b) What change is there in the entropy of the universe due to the melting of the ice? (specific heat of water is $4218 \text{ J K}^{-1} \text{ kg}^{-1}$)

思考题

By differentiating the enthalpy function ($h = u + p\alpha$), show that

$$\left(\frac{\partial p}{\partial T}\right)_s = \left(\frac{\partial s}{\partial \alpha}\right)_p$$

where s is entropy.

Note:

$$dh = Tds + \alpha dp$$

$$\frac{\partial}{\partial x_j} \left(\frac{\partial y}{\partial x_i} \right) = \frac{\partial}{\partial x_i} \left(\frac{\partial y}{\partial x_j} \right)$$

Show that this is equivalent to the Clausius-Clapeyron Equation.

This is one of the Maxwell's four thermodynamic equations.

Radiative Equilibrium Temperature: Vertical Structure

Radiative equilibrium for stratified atmosphere:

$$\rho \frac{\delta q}{\delta t} = - \frac{\partial F}{\partial z} = 0$$

↪

$$F = F_T - F_S = \text{constant} = 0$$

Assuming **no extinction of solar radiation by atmosphere**

So, solar radiation flux (downward is negative):

$$F_S = \mu_0 \cdot S \cdot (1 - R) = \text{constant}$$

$$\mu_0 = 1/4 \quad \text{for global mean annual mean}$$

$$\mu_0 = 1/\pi \quad \text{for tropical mean annual mean}$$

Thus: $F_T = \text{constant}$

Radiative Equilibrium Temperature: Vertical Structure

For radiative transfer of thermal radiation flux for **stratified atmosphere with no scattering**:

$$\frac{dF_T}{d\tau} = 4\pi(1 - \omega_{0_T})(\bar{I} - B) = 0$$

$$\frac{d^2F_T}{d\tau^2} = 3(1 - \omega_{0_T})(1 - \omega_{0_Tg})F_T - 4\pi(1 - \omega_{0_T})\frac{dB}{d\tau} = 0$$

$$\omega_{0_T} = 0$$

We get: $\bar{I} = B$

$$4\pi \cdot \frac{dB}{d\tau} = 3F_T$$
$$B(\tau) - B(0) = \frac{3F_T\tau}{4\pi}$$

Radiative Equilibrium Temperature: Vertical Structure

Solution:

$$B = \bar{I}$$
$$B(\tau) - B(0) = \frac{3F_T\tau}{4\pi}$$

Boundary conditions:

$$B_g - B(\tau_1) = \frac{F_T}{2\pi}$$
$$B(0) = \frac{F_T}{2\pi}$$

Thus:

$$B(\tau) = \frac{\sigma}{\pi} T(\tau)^4 = \frac{F_T}{2\pi} \left(1 + \frac{3\tau}{2} \right)$$

$$B(\tau_1) = \frac{\sigma}{\pi} T(\tau_1)^4 = \frac{F_T}{2\pi} \left(1 + \frac{3\tau_1}{2} \right)$$

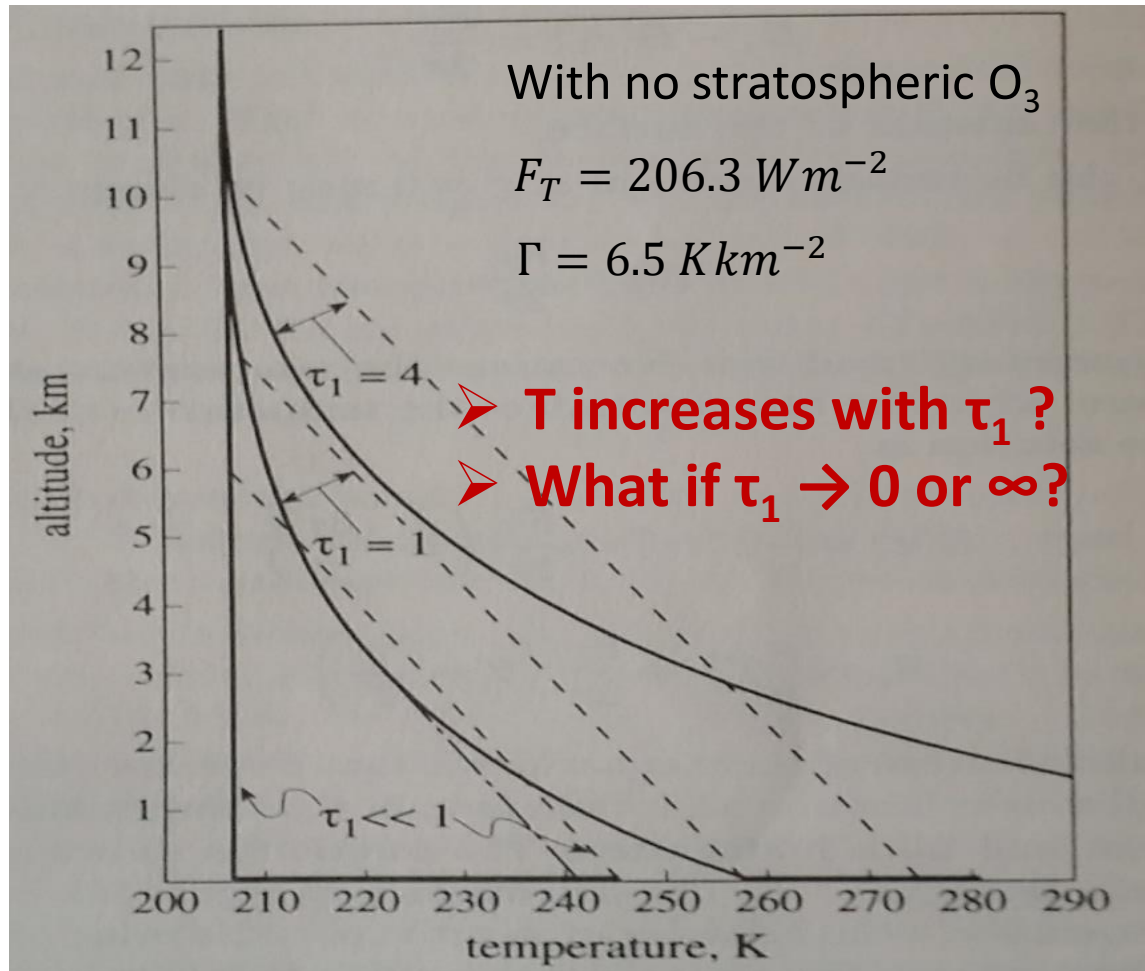
$$B_g = \frac{\sigma}{\pi} T_g^4 = \frac{F_T}{2\pi} \left(2 + \frac{3\tau_1}{2} \right)$$

Temperature
Discontinuity
at τ_1

Temperature Discontinuity

In many cases (e.g., CO₂): $\tau(z) = \tau_1 \cdot e^{(-z/H)}$ H: scale height

Thus:



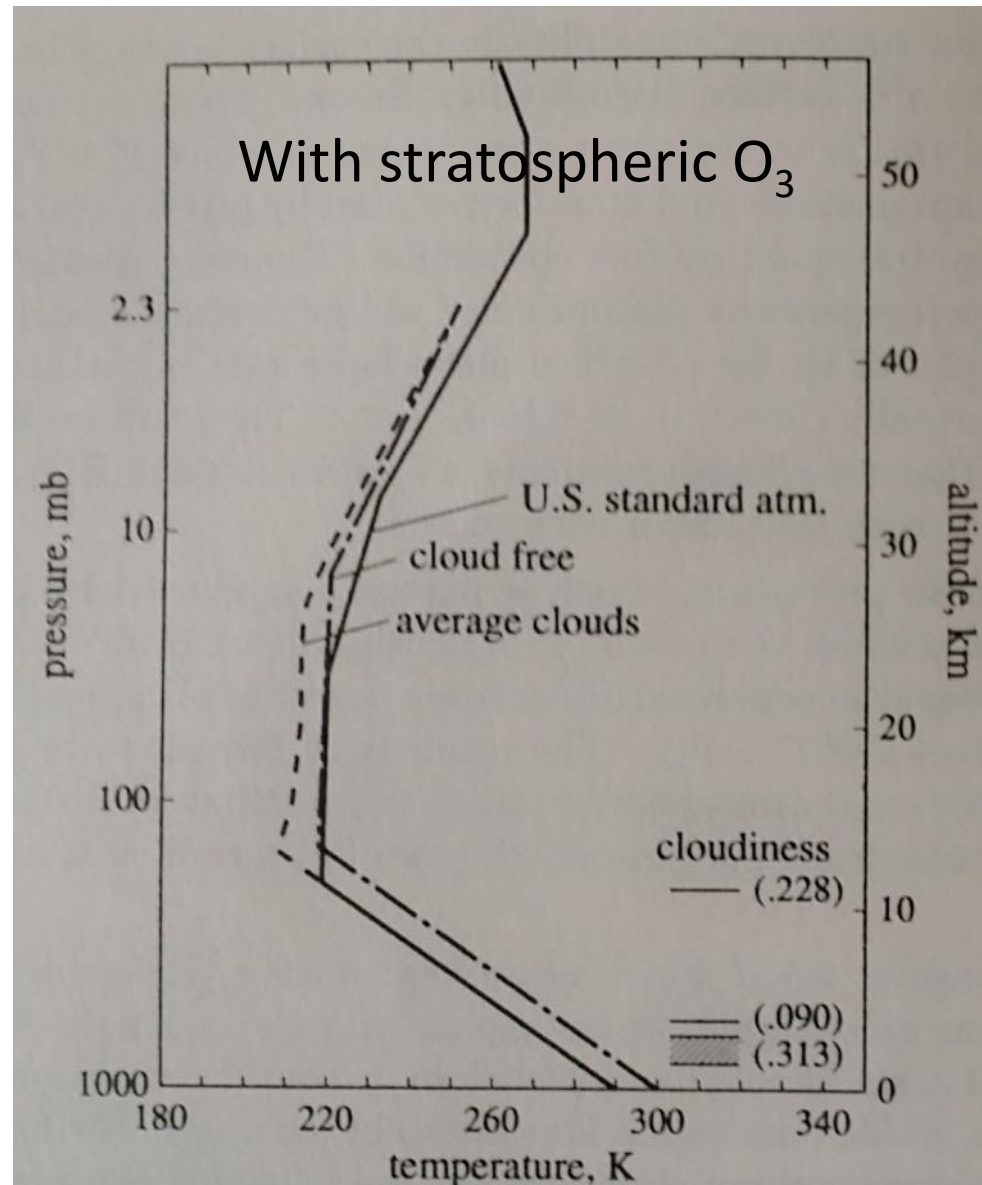
Radiation-Convection in the Troposphere

$$F_{s_cloudfree} = 299 \text{ Wm}^{-2}$$

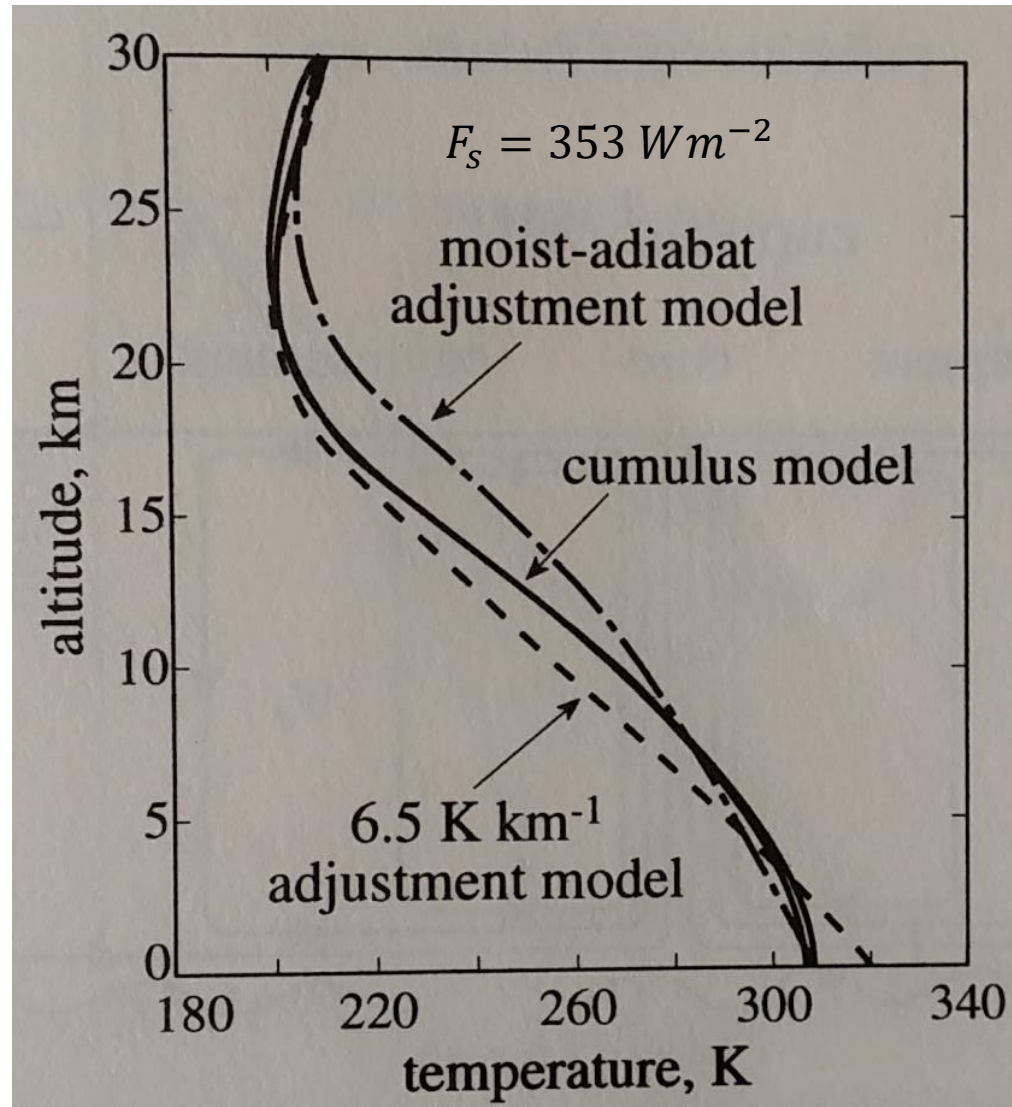
$$F_{s_cloudy} = 228 \text{ Wm}^{-2}$$

$$\Gamma = 6.5 \text{ Kkm}^{-1}$$

$$\text{SZA} = 60^\circ$$



Radiation-Convection in the Tropical Troposphere



The Emission Level

At the emission level τ_e ,
black body emission is equal to planetary emission

$$F_T(\tau = 0) = \pi B_e = \sigma T_e^4$$

$$B_e = B(\tau_e) = \frac{F_T}{2\pi} \left(1 + \frac{3\tau_e}{2} \right) = \frac{F_T}{\pi}$$

Therefore:

$$\tau_e = \frac{2}{3}$$

The Chapman Layer

From the thermal radiance equation (**assuming $\tau_1 \gg 1$**):

$$\begin{aligned}
 I_T^+(\tau = 0, \mu) &= B(T_g) \cdot e^{-\tau_1/\mu} + \frac{1}{\mu} \cdot \int_0^{\tau_1} B(T(\tau)) \cdot e^{-\tau/\mu} \cdot d\tau && \text{at TOA} \\
 &\approx \frac{1}{\mu} \cdot \int_0^{\tau_1} B(T(\tau)) \cdot e^{-\tau/\mu} \cdot d\tau \\
 &= \int_0^{\infty} B(z) \frac{-1}{\mu} \cdot \frac{d\tau}{dz} \cdot e^{-\tau/\mu} \cdot dz = \int_0^{\infty} B(z) \cdot h(z, \mu) \cdot dz
 \end{aligned}$$

Where:

$$\begin{aligned}
 h(z, \mu) &= \frac{-1}{\mu} \cdot \frac{d\tau}{dz} \cdot e^{-\tau/\mu} && \text{Kernel function, i.e., } \int_0^{\infty} h(z, \mu) dz = 1 \\
 &= \frac{1}{H\mu} \cdot \tau(z) \cdot e^{-\tau(z)/\mu} && [\text{for } \tau(z) = \tau_1 \cdot e^{-z/H}]
 \end{aligned}$$

At maximum h : $h(max) = \frac{1}{eH}$ $\tau(max) = \mu$

The Chapman Layer

Mean value theorem for thermal radiation flux:

$$\begin{aligned} F_T^+(z) &= \int_0^{2\pi} \int_0^1 I_T^+(z, \mu) \cdot \mu \cdot d\mu \cdot d\varphi \\ &= 2\pi \cdot \int_0^1 I_T^+(z, \mu) \cdot \mu \cdot d\mu \\ &= \pi \cdot I_T^+(z, \hat{\mu}) \end{aligned}$$

Many investigators found that globally, $\hat{\mu} = \frac{2}{3}$

Therefore,
the kernel function for flux to space peaks at the emission level

The Chapman Layer

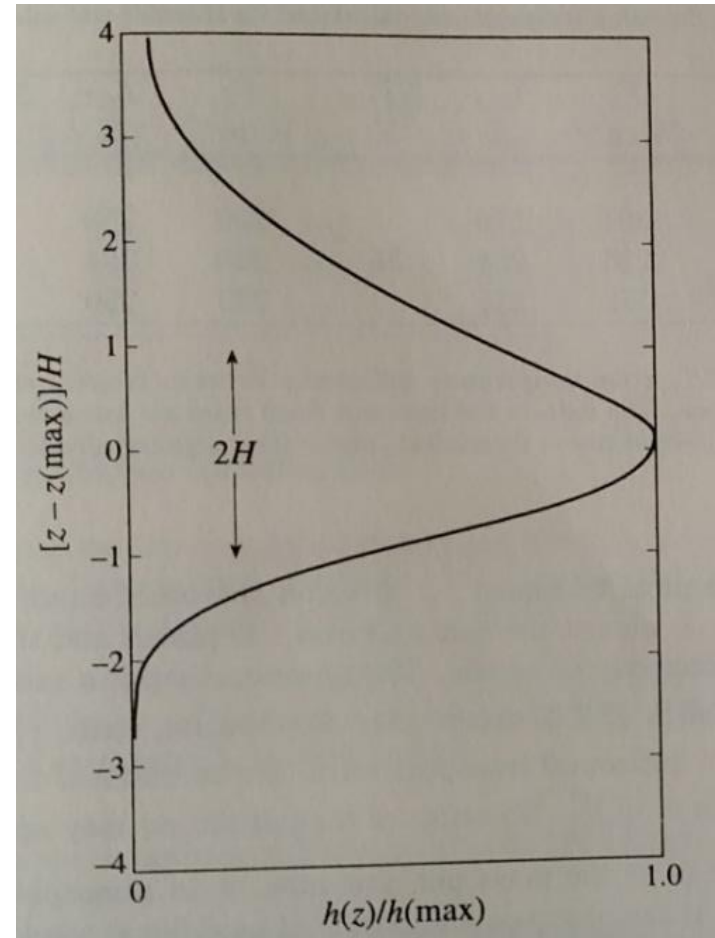
Re-arrange the kernel function:

$$\frac{h(z)}{h(max)} = \frac{\tau(z)}{\tau(max)} e^{1 - \frac{\tau(z)}{\tau(max)}}$$

Where:

$$h(max) = \frac{1}{eH}$$

$$\frac{\tau(z)}{\tau(max)} = e^{-\frac{z-z(max)}{H}}$$



Semi-gray Greenhouse Model

Semi-gray model:

$$\tau_e = \tau_1 e^{-z_e/H} = 2/3$$

$$\rightarrow z_e = H \ln \frac{3\tau_1}{2}$$

$$\rightarrow T_g = T_e + \Gamma_{obs} H \ln \frac{3\tau_1}{2}$$

Given that:

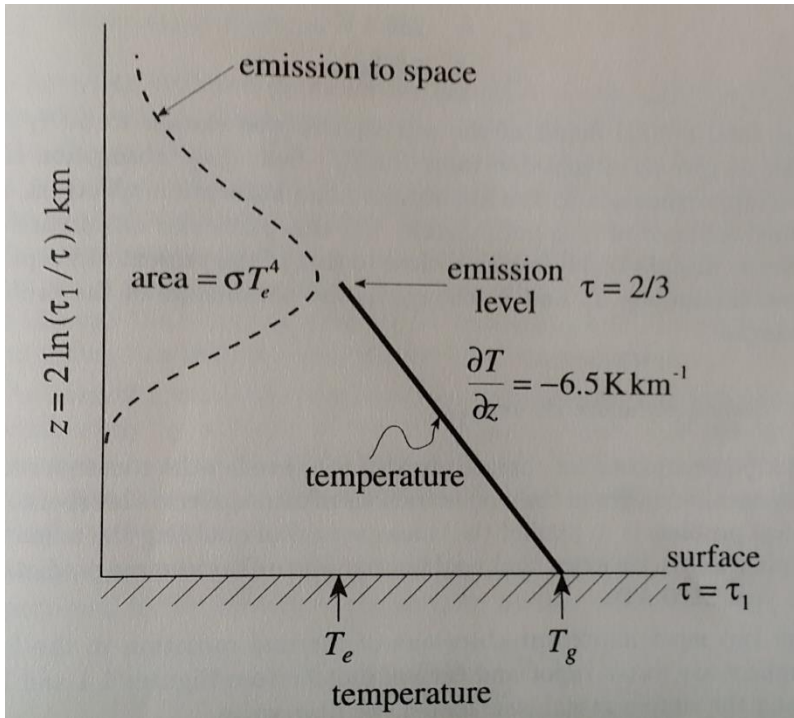
$$F_S = 236.6 \text{ W m}^{-2}$$

$$\Gamma_{obs} = 6.5 \text{ K km}^{-1}$$

$$H = 2 \text{ km}$$

$$\tau_1 = 8$$

} For H₂O



Therefore:

$$z_e = 5.0 \text{ km}$$

$$T_e = 254.1 \text{ K}$$

$$T_g = 286.4 \text{ K}$$

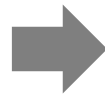
Climate Sensitivity with No Feedbacks other than Planck's

$$\frac{\partial T_g}{\partial \ln \tau_{1,CO_2}} = \Gamma_{obs} H \frac{\partial \ln \tau_1}{\partial \ln \tau_{1,CO_2}} = \Gamma_{obs} H \frac{\tau_{1,CO_2}}{\tau_1}$$

Current atmosphere,
CO₂ and H₂O only:

$$\tau_1 = \tau_{1,CO_2} + \tau_{1,H_2O}$$

$$\frac{\tau_{1,CO_2}}{\tau_1} \approx 0.125$$



Climate sensitivity is small:

$$\frac{\partial T_g}{\partial \ln \tau_{1,CO_2}} = 1.6 \text{ K}$$

$$\Delta T_g = 1.1 \text{ K} \quad \text{for } 2 \times [\text{CO}_2]$$

Water Vapor Feedback

Clausius-Clapeyron Equation

At constant
air pressure

$$\frac{1}{w_s} \frac{dw_s}{dT} \approx \frac{1}{e_s} \frac{de_s}{dT} \approx \frac{L_v}{R_v T^2}$$

Equilibrium (saturation) water vapor mixing ratio near the surface will increase by 7.2% for every 1 K increase in temperature around 273 K, that is,

w_s will double for every 10 K increase in T

$$w_s = w_{s,0} \cdot e^{\frac{T_g - T_{g,0}}{10} \ln 2}$$

Climate Sensitivity with Water Vapor Feedback

If surface RH and vertical profile of water vapor are unchanged:

$$\tau_{1,H_2O} = \tau_{1,H_2O,0} \cdot e^{\frac{T_g - T_{g,0}}{10} \ln 2}$$

We get:

$$\frac{\partial T_g}{\partial \ln \tau_{1,CO_2}} = \alpha \cdot \Gamma_{obs} H \frac{\tau_{1,CO_2}}{\tau_1}$$

$$\Delta T_g = 5.2 K \text{ for } 2 \times [CO_2]$$

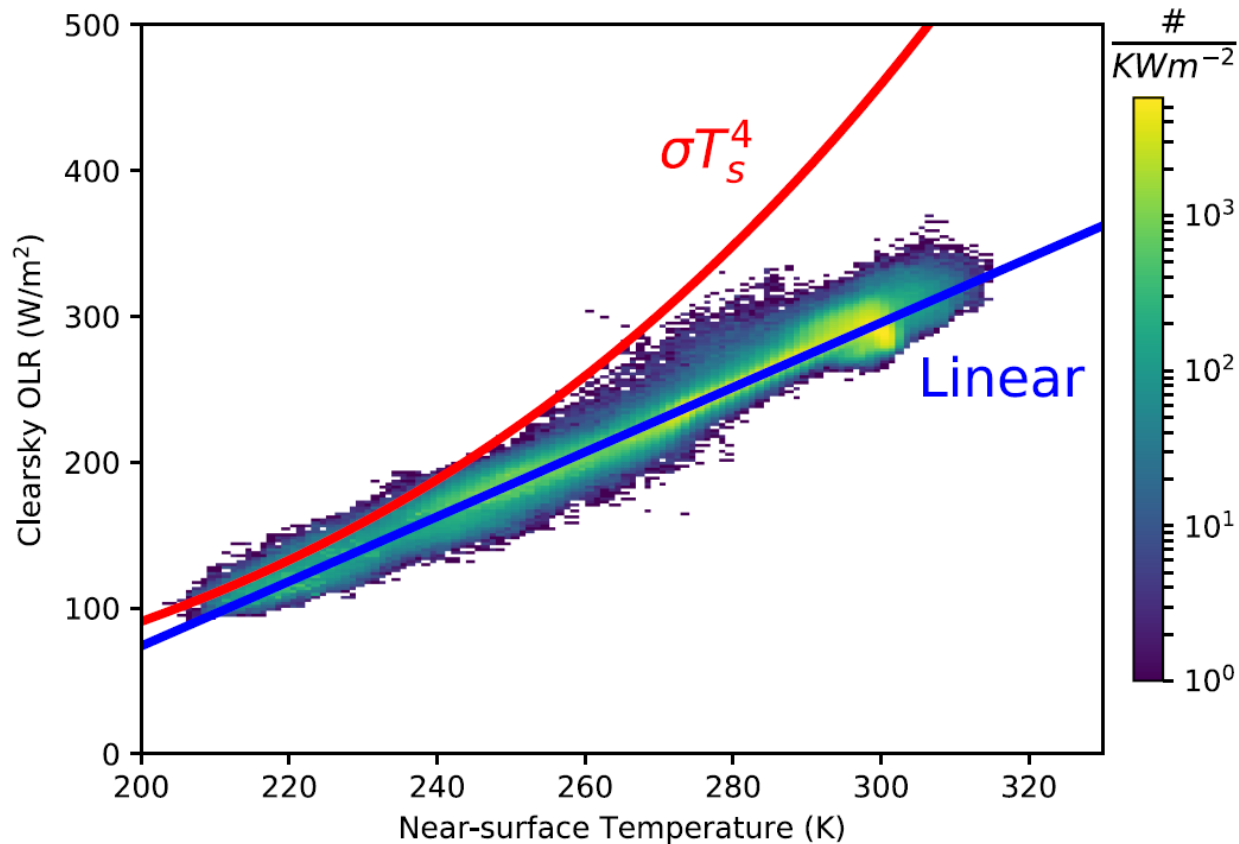
- **H₂O feedback greatly enhances climate sensitivity**
- **IPCC AR6 estimate: 2.5-4.0 K (likely range)**
- **Runaway greenhouse effect?**

Where:

$$\alpha = \left(1 - \Gamma_{obs} H \frac{1}{\tau_1} \frac{\partial \tau_{1,H_2O}}{\partial T_g} \right)^{-1} = \left(1 - \Gamma_{obs} H \frac{\tau_{1,H_2O} \ln 2}{\tau_1 10} \right)^{-1} = 4.7$$

Linking OLR and Near-Surface Temperature

$$T_e = T_g - \Gamma_{obs} H \ln \frac{3\tau_1}{2}$$



Effects of Clouds

- Low clouds (≤ 4 km): changes albedo
- High clouds (~ 8 km): changes albedo and emission layer height

	No cloud			Complete cloud			Diff.
	R	z_e , km	T_g , K	R	z_e , km	T_g , K	ΔT_g , K
High clouds	0.12	5	302	0.3	8	307	+5
Low clouds	0.12	5	302	0.7	5	234	-68

Energy Balance Climate Models: Meridional Change

$$C \frac{\partial T_g(x)}{\partial t} = \text{solar term} + \text{thermal term} + \text{dynamic term}$$

$$\text{solar term} = \frac{S}{4} \cdot s(x) \cdot [1 - A(x)]$$

$$s(x) = 1 - 0.241(3x^2 - 1)$$

$$\text{thermal term} = -[211.1 + 1.55T_g(x)]$$

$$\text{dynamical term} = \frac{\partial}{\partial x} \left[(1 - x^2) \cdot D \cdot \frac{\partial T_g(x)}{\partial x} \cdot s(x) \right]$$

S = Solar irradiance

x = sin (latitude)

D = diffusion coefficient

A = Surface albedo

T_g = Ground temp. in °C

Energy Balance Climate Models

Albedo:

- Ice or snow: $A_i = 0.6, T_g \leq -10^\circ\text{C}$
- Ocean or land surface: $A_s = 0.3, T_g > -10^\circ\text{C}$

Steady state ($\frac{\partial T_g}{\partial t} = 0$), and no dynamical term ($D = 0$):

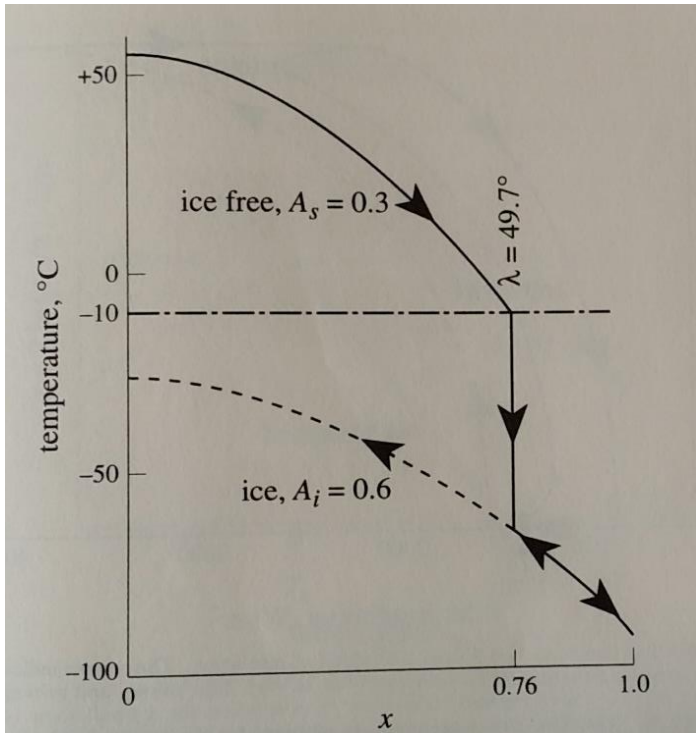
$$T_g(x) = \frac{(1 - A) \cdot S}{4 \times 1.55} [1 - 0.241(3x^2 - 1)] - \frac{211.1}{1.55}$$

$$S = \frac{(211.1 - 10 \times 1.55) \times 4}{(1 - A) \cdot [1 - 0.241(3x_i^2 - 1)]}$$

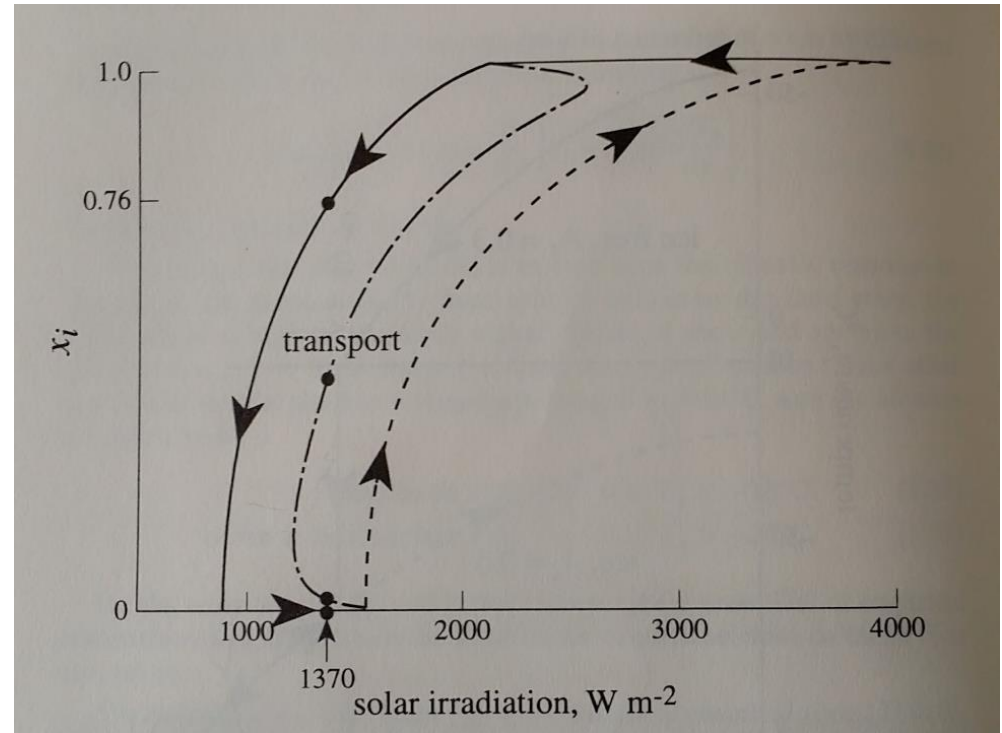
x_i represents ice/snow line (where $T_g = -10^\circ\text{C}$)

Energy Balance Climate Models

Temperature as a function of latitude



Position of the ice line as a function of insolation



Hysteresis 迟滞现象:

The dependence of the output of a system not only on its current input, but also on its history of past inputs

思考题

- ✓ For a planet similar to the Earth but with no water, how would its surface temperature and air temperature be like? Consider a two-box model under energy balance.
- ✓ Why is there a lapse rate feedback? What are the causes of the meridional dependence of this feedback?
- ✓ How would surface temperature, air temperature and water vapor change if anthropogenic greenhouse gas concentrations continue to increase? Runaway greenhouse effect?